

Spectroscopic Monitoring of 10 Northern SPB Candidates¹

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Abstract. We present the results of 1.5 years of spectroscopic monitoring of 10 bright northern SPB candidates discovered from the Hipparcos mission. All stars have line-profile variations and are periodic, with periods corresponding to high-order g -modes. We therefore confirm them to be new SPBs.

1. Introduction

Because of the many awaited space missions, the coming decade will be important for asteroseismology. In order to study the deep internal layers, g -modes pulsators, such as Slowly Pulsating B stars (hereafter SPBs), are the best objects. The HIPPARCOS mission led to the discovery of many new SPBs candidates (Waelkens et al., 1998). However, some of these candidates turned out not to be SPBs (see De Cat, 2001), and spectroscopic data are needed to confirm the pulsational nature of the targets.

A total of 10 objects was selected from the list provided by Waelkens et al. (1998). Note that HD 147394 is discussed separately (Briquet et al., these proceedings). Depending on the brightness of each star, between 15 and 60 spectra were obtained, with a signal-to-noise ratio around 100 and a resolving power of 15 000 in the Si II $\lambda\lambda$ 4128, 4130 region.

¹The follow-up data were obtained with the 1.52m telescope of the Haute-Provence Observatory in France

2. Results

The frequency analysis was performed both on HIPPARCOS (search for multi-periodicity) and radial velocity data (validation of the Hipparcos frequency). Results are provided in Table 1.

Table 1. Pulsation characteristics of the SPBs. The f_p and f_s values, respectively, correspond to the dominant photometric and spectroscopic frequencies. The pulsation constant Q_p is computed from the f_p value. No value f_s is provided for HD 140873 because it is a double-lined spectroscopic binary and the companion could not be removed.

Star (HD)	f_p [cd ⁻¹]	f_s [cd ⁻¹]	Q_p [cd ⁻¹]
1976	0.93914	0.39934	0.38
21071	1.18843	1.18843	0.52
25558	0.65284	0.65284	0.69
28114	0.79104	0.79104	0.42
138764	0.794	0.794	0.62
140873	1.149	-	0.55
182255	0.7922	0.7922	0.67
206540	0.7624	0.7624	0.54
208057	0.80172	0.89045	0.48

All target stars show line profile variations on time scales expected for non-radial g -mode pulsations. All but one, HD 28114, show evidence for multi-periodicity. The latter star still should be considered as an SPB star since line profiles are well-marked, the radial velocity curve is sinusoidal and the star presents “normal” abundances (Adelman & Philip, 1996). See also Briquet et al. (these proceedings).

Two stars (HD 1976 and HD 208057) present different dominant frequencies in spectroscopy and photometry: this could indicate different modes of high and low degree, respectively, but both stars are also rapid rotators and possibly members of binary systems.

References

- Adelman, S.J. & Philip, A.G.D. 1996, MNRAS, 282, 1181
 De Cat, P. 2001, PhD Thesis, University of Leuven, Belgium
 Waelkens, C., Aerts, C., Kestens, E., Grenon, M., & Eyer, L. 1998, A&A, 330, 215