

Evidence of a Low-Mass Companion to AB Doradus

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Abstract. From the combination of VLBI phase-referenced observations and Hipparcos satellite data, we have found evidence of a low-mass object orbiting the late-type star AB Doradus. The mass of the new object is near the hydrogen burning limit and will constitute a precise point for calibrating the low end of the main sequence. This represents the first detection of a low-mass stellar companion using the VLBI technique, which could become an important tool in future searches for planets and brown dwarfs orbiting other stars.

1. Introduction

Determination of star's kinematics with submilliarcsecond precision provides not only the position, proper motion, and parallax of the observed star but possible accelerations caused by the gravitational interaction with companion objects. Unlike radial velocity techniques, astrometry-based techniques, such as VLBI or Hipparcos satellite data, are able to determine a full set of orbital elements which provides an unambiguous estimate of the mass of the companion.

2. Observations and Results

We have combined two astrometric data types, VLBI phase-referenced data and Hipparcos satellite data, to monitor the position of the Southern K0-star AB Doradus (=HD 36705 = HIP 25647, AB Dor hereafter). The 8.4 GHz VLBI data were originally intended to contribute to the determination of the link between the Hipparcos and VLBI reference frames (Lindegren & Kovalevsky 1995). The VLBI array was composed of the Australian antennas at Tidbinbilla, Parkes, and Hobart; the data were correlated at the MkIII A correlator of the US Naval Observatory at Washington, DC. We followed the astrometric VLBI technique described in Lestrade et al. (1990) to estimate the star's coordinates, referred to the IERS reference frame. The coordinates from each data type, VLBI and Hipparcos satellite, showed independent evidence of a previously unseen non-linear motion on the sky of AB Dor, indicating that this star was likely an astrometric double. Since the final Hipparcos catalogue (ESA 1997) is also linked to the IERS reference system (Kovalevsky et al. 1997), we could use the joint VLBI

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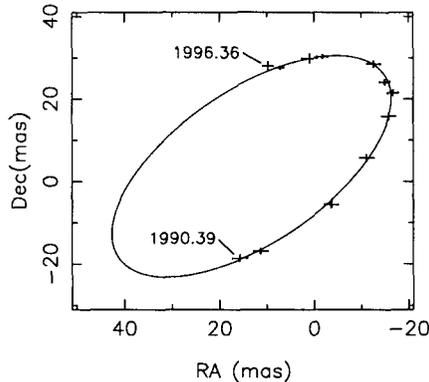


Figure 1. The apparent orbit for the reflex motion of AB Dor corresponding to a companion mass of $0.094 M_{\odot}$. The five earliest epochs correspond to Hipparcos data. The continuous line represents the least-square fitted reflex orbital motion of AB Dor.

+ Hipparcos data set to estimate simultaneously the position, proper motion, parallax, and orbital elements of the reflex motion of AB Dor. Although the combined data did not cover a full orbit, we investigated the astrometric and orbital parameter space to find orbits compatible with our positions (see Guirado et al. 1997 for a detailed analysis). The main results of our search are: *i*) a new value of the parallax of AB Dor ($66.3 < \pi < 67.2$ milliarcseconds), which, in turn, provided a revised mass for AB Dor, $0.76 M_{\odot}$, and *ii*) the dynamical mass of the unseen companion, which is constrained to lie between 0.08 and $0.11 M_{\odot}$ (see Fig 1).

The newly discovered companion, AB Dor C, is one of the few low mass objects near the hydrogen burning limit whose mass has been determined dynamically (Henry & McCarthy 1993). A suitable photometric calibration would locate this object in a mass-luminosity diagram and would constitute a new precise point for the calibration of the low end of the main sequence (e.g., Baraffe et al. 1995). The detection of AB Dor C demonstrates the capability of VLBI phase-referencing, enhanced in this case with Hipparcos data, for detecting very low-mass companions orbiting stars.

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