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The results from a search for high-velocity hydrogen around the direction to the galactic center are presented. About 2000 positions were surveyed with the 43-m radiotelescope of NRAO with a rms of 0.03 K on a velocity interval of -1000 to +1000 km s $^{-1}$.

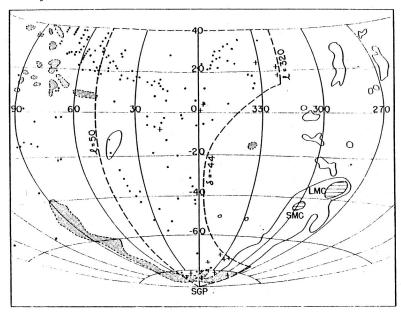


Figure 1. Distribution of HI with $|\rm V_{LSR}|>80~km~s^{-1}$ in the inner Galaxy. The area surveyed by us is limited by the broken lines at $\rm \pounds =50^{\circ}$, b=+40° and $\rm \delta = -44^{\circ}$, the southern limit of the observable sky with the 43-m telescope of NRAO. Data outside that area are from surveys by other authors (Mirabel, 1981 and references therein). Dots are clouds smaller than 3° with V_{LSR} $< -80~km~s^{-1}$. Crosses are clouds smaller than 3° with V_{LSR} $> +80~km~s^{-1}$. Contours with hatching are clouds greater than 3° with V_{LSR} $< -80~km~s^{-1}$. Contours without hatching are clouds greater than 3° with V_{LSR} $> +80~km~s^{-1}$.

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H. van Woerden et al. (eds.), The Milky Way Galaxy, 411-412. © 1985 by the IAU.

Figure 1 shows the presence of several streams or complexes. In the first place, there appears a stream that extends from $\ell=5^{\circ}$ b=+5° to $\ell=90^{\circ}$ b=+40°. This stream has velocities in the range of -80 to -140 km s⁻¹ and extends up to $\ell=120^{\circ}$ b=+55° (Giovanelli, 1980). On the other hand, there seems to be a stream of small and faint clouds with $V_{LSR} <$ -140 km s⁻¹ extending from $\ell=20^{\circ}$ b=-55° to $\ell=48^{\circ}$ b=+26°. It is interesting that this stream could stretch out from the southern into the northern galactic hemisphere, across the galactic plane.

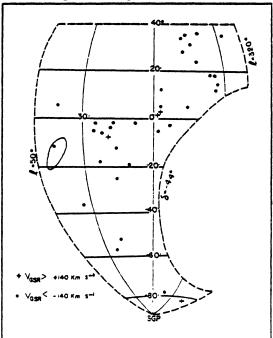


Figure 2. Distribution of clouds with very high galactocentric velocities ($|V_{\rm GSR}| > 140~{\rm km~s^{-1}}$) in the region surveyed around the direction to the galactic center. More than 85% of the very-high-velocity detections are negative in sign. There appears a complex of clouds with very high negative velocities in the region $330^{\circ} < 1 < 30^{\circ} - 40^{\circ} < 1 < 30^{\circ} < 1 < 1 < 30^{\circ} < 1 < 30^{\circ}$

Since similar extreme velocities of about $-200~\rm km~s^{-1}$ are observed in the galactic center and anticenter, these results are strong evidence for a high-velocity inflow of neutral hydrogen toward the Milky Way. Assuming that the gas that is infalling with velocities greater than 140 km s⁻¹ is at a distance of 20 kpc from the galactic center, we estimate a present net influx of 0.2 $\rm M_{\odot}$ yr⁻¹ toward the Galaxy.

This research was supported by NSF grant AST 80-13148.

REFERENCES

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