SIZE OF THE BROAD LINE REGION IN NGC 3227: RESULTS FROM THE LAG CAMPAIGN *

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We present the results of a monitoring campaign of the AGN in NGC 3227, conducted at La Palma by the LAG european consortium. The data are long slit spectra, obtained with the 1.5m INT and 4.2m WHT, from January to June 1990. The slit width was 1.5 arcsec. The spectral resolution was about 3 A.

From the observed MgI λ 5167, 5173, 5183 in the NGC 3227 spectra we deduce that the stellar contribution is at most 40% at λ 5180 A when the AGN is in his lower state.

To separate the BLR and NLR emissions we have performed a gaussian decomposition of the blends $H\alpha + [NII] + [SII]$ and $H\beta + [OIII]$. We have constrained as many parameters as possible, following physical arguments. The same blend modeling was used for all epochs, only the amplitude of the BLR components was allowed to vary.

After having performed the line blend decomposition, subtracted the stellar population contribution and applied corrections for seeing differences, we have obtained the light curves for the BLR H α and H β line emission as well as for the visible continuum (derived under the assumption that the NLR emission does not vary with time). We find out variation amplitudes around 40% over timescales of 6 weeks. The cross correlation of the line and continuum light curves provides a BLR size of 17 +/- 7 days.

The Balmer line difference between the highest and lowest levels shows a symmetrical flat-topped profile with FWHM ≈ 2900 km/s. Assuming a keplerian broadening of the lines, and that the accretion disc is coplanar with the galactic disc, we conclude that the bulk of the BLR emission occurs at 10^3 Schwarzschild radii and that the central object has a mass of $2 \times 10^8 M_o$.

From a small sample of AGN (7) we find that the BLR size is proportional to L^{γ} , where $\gamma < 0.5$, instead of 0.5 assumed in the models. A correlation analysis indicates a slope of about 0.2.

References

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* The Lovers of Active Galaxies (LAG) consortium was set up to apply for international time on the Canaries Observatories and conducted a series of programs related to the physics of AGN.

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