

Strong magnetic spiral pattern in a ringed galaxy NGC 4736

Krzysztof T. Chyży¹ and Ronald Buta²

¹Astronomical Observatory of the Jagiellonian University, ul. Orla 171, 30-244 Kraków, Poland

²Department of Physics and Astronomy, University of Alabama, USA

Abstract. We present high sensitivity radio polarimetric (VLA) observations of a galaxy with strong orbital resonances – NGC 4736. The total radio intensity at 8.4 GHz covers smoothly the whole galaxy bulge and reveals a distinct ring of radio emission closely related to the ring morphology visible in infrared, CO and H α emission. However, the magnetic field reveals a very coherent spiral pattern. The magnetic field vectors are crossing the inner starbursting ring, the dust lanes within the ring and other rather circularly shaped features visible in other gas traces. Either the magnetic field uncovers the pattern of gas motions not seen in other spectral ranges, or the spiral magnetic field is of a pure dynamo origin, ignoring the ringed morphology of the galaxy.

Keywords. galaxies: ISM, galaxies: spiral, individual (NGC4736), ISM: magnetic fields

1. Radio polarimetric observations of a ringed galaxy

How the galactic magnetic field can operate in early Hubble-type galaxies and in the presence of strong orbital resonances is as yet unexplored. Using the VLA we investigated this problem and observed NGC 4736 (M 94) – the nearest and largest galaxy showing a ringed morphology (Buta & Combes 1996). Its inner and outer rings are believed to coincide with Lindblad resonances of a triaxial bulge, the inner disk (oval distortion) and/or the nuclear minibar. The giant inner ring located at a distance of about 47'' (1.1 kpc) from the (LINER) nucleus, is a region of intense star formation which dominates the UV, H α , infrared and CO images (Wong & Blitz 2001).

We observed NGC 4736 with the VLA in its D array at 8.4 and 4.6 GHz. The observed total radio intensity covers smoothly the whole galaxy bulge and reveals a distinct ring of radio emission (Fig. 1). The radio ring is closely related to the ring morphology visible in other spectral ranges, but the radio polarized emission at 8.4 GHz seems to be not clearly associated with H α emitting gas, or other ISM phases. Instead, it reveals a very coherent spiral pattern of the regular magnetic field (Fig. 1).

Surprisingly, the prominent inner ring does not tune the magnetic field vectors. Two magnetic spiral arms of constant pitch angle of about 45° emerge from the centre of NGC 4736 and cross the inner ring. This is opposite to what is observed in grand-design spiral galaxies where the regular magnetic field typically follows the shape of nearby gaseous spiral arms, even if a galaxy is perturbed (Chyży *et al.* 2007).

Between the arms, and inside the ring, the regular field of NGC 4736 has a smaller pitch angle. This seems to be related to a central minibar extending along the galaxy's minor axis. The magnetic field vectors also cross the NW long and strong dust lane, visible in optical images. A spiral network of dust armlets in the central part of the galaxy (probably of acoustic origin) seems also to be not associated with the magnetic spiral – the magnetic vectors have typically larger pitch angles, by at least 20°. Nevertheless, the

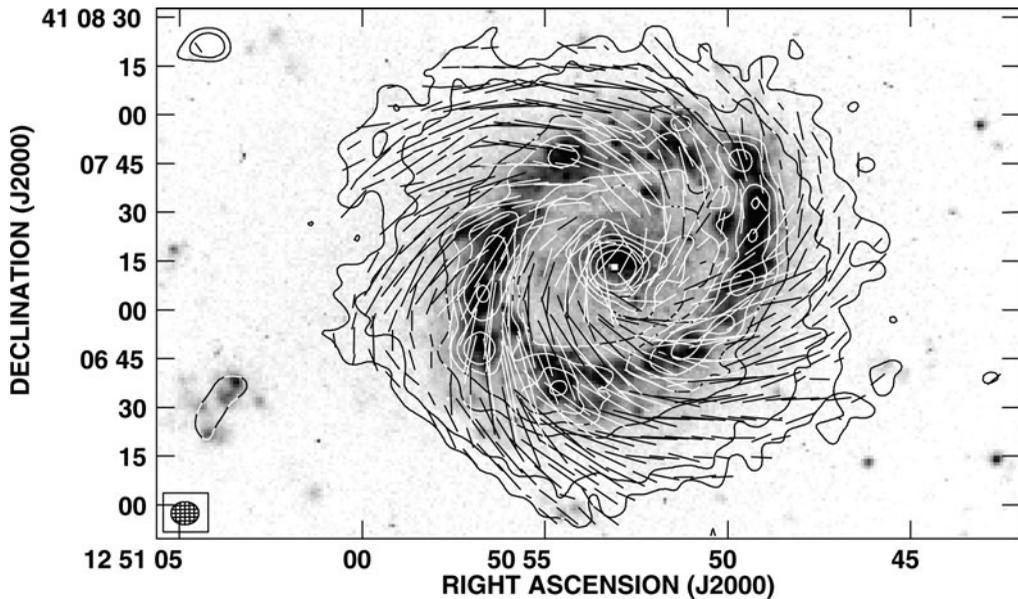


Figure 1. The total radio intensity contour map of NGC 4736 at 8.4 GHz at $8'' \times 7''$ resolution with observed magnetic field vectors of the polarized emission overlaid on the $H\alpha$ image (from Knapen *et al.* 2003). Contours are at (3, 8, 32, 50, 80, 128, 256) $\times 6 \mu\text{G}/\text{b.a.}$. Vectors of length of $10''$ correspond to polarized intensity of $26 \mu\text{G}/\text{b.a.}$.

magnetic field may still help in channelling the surrounding gas and dust into an active LINER nucleus.

2. Magnetic spiral

The discovered spiral structure of the magnetic field in NGC 4736 seems to contradict the galaxy's ringed morphology. If the magnetic field is totally frozen into the interstellar plasma then the observed magnetic vectors may reveal the pattern of the galactic gas motions and shocks which is not yet discerned in other spectral ranges. On the other hand, and more likely, the spiral magnetic field can be of pure dynamo origin. In this case the magnetic field and the interstellar turbulence must be strong enough to resist the local gas flows and cross the star-forming ring imperceptibly. In each case NGC 4736 seems to be a key object to study the resonant dynamics and a dynamo process of the magnetic field generation without spiral geometry of density waves.

Acknowledgements

This work was supported by the Polish Ministry of Science and Higher Education, grant 2693/H03/2006/31.

References

- Chyży, K. T., Ehle, M., & Beck, R. 2007, AA, in press, astro-ph 0708.1533
- Buta, R. & Combes, F. 1996, *Fund. Cosmic Phys.* 17, 95
- Wong, T. & Blitz, L. 2001, *ApJ* 540, 771
- Knapen, J. H., de Jong, R. S., Stedman, S., & Bramich, D. M., 2003, *MNRAS*, 344, 527